

FHSST Authors

# The Free High School Science Texts: Textbooks for High School Students Studying the Sciences Chemistry Grades 10 - 12

Version 0 November 9, 2008 Copyright 2007 "Free High School Science Texts"

Permission **is** granted to copy, distribute and/or modify this document under the terms of the GNU Free Documentation License, Version 1.2 or any later version published by the Free Software Foundation; with no Invariant Sections, no Front-Cover Texts, and no Back-Cover Texts. A copy of the license is included in the section entitled "GNU Free Documentation License".



Did you notice the **FREEDOMS** we've granted you?

Our copyright license is **different!** It grants freedoms rather than just imposing restrictions like all those other textbooks you probably own or use.

- We know people copy textbooks illegally but we would LOVE it if you copied our's go ahead copy to your hearts content, **legally!**
- Publishers' revenue is generated by controlling the market, we don't want any money, go ahead, distribute our books far and wide we DARE you!
- Ever wanted to change your textbook? Of course you have! Go ahead, change ours, make your own version, get your friends together, rip it apart and put it back together the way you like it. That's what we really want!
- Copy, modify, adapt, enhance, share, critique, adore, and contextualise. Do it all, do it with your colleagues, your friends, or alone but get involved! Together we can overcome the challenges our complex and diverse country presents.
- So what is the catch? The only thing you can't do is take this book, make a few changes and then tell others that they can't do the same with your changes. It's share and share-alike and we know you'll agree that is only fair.
- These books were written by volunteers who want to help support education, who want the facts to be freely available for teachers to copy, adapt and re-use. Thousands of hours went into making them and they are a gift to everyone in the education community.

## FHSST Core Team

Mark Horner ; Samuel Halliday ; Sarah Blyth ; Rory Adams ; Spencer Wheaton

## FHSST Editors

Jaynie Padayachee ; Joanne Boulle ; Diana Mulcahy ; Annette Nell ; René Toerien ; Donovan Whitfield

## FHSST Contributors

Rory Adams ; Prashant Arora ; Richard Baxter ; Dr. Sarah Blyth ; Sebastian Bodenstein ; Graeme Broster ; Richard Case ; Brett Cocks ; Tim Crombie ; Dr. Anne Dabrowski ; Laura Daniels ; Sean Dobbs ; Fernando Durrell ; Dr. Dan Dwyer ; Frans van Eeden ; Giovanni Franzoni ; Ingrid von Glehn ; Tamara von Glehn ; Lindsay Glesener ; Dr. Vanessa Godfrey ; Dr. Johan Gonzalez ; Hemant Gopal ; Umeshree Govender ; Heather Gray ; Lynn Greeff ; Dr. Tom Gutierrez ; Brooke Haag ; Kate Hadley ; Dr. Sam Halliday ; Asheena Hanuman ; Neil Hart ; Nicholas Hatcher; Dr. Mark Horner; Robert Hovden; Mfandaidza Hove; Jennifer Hsieh; Clare Johnson ; Luke Jordan ; Tana Joseph ; Dr. Jennifer Klay ; Lara Kruger ; Sihle Kubheka ; Andrew Kubik ; Dr. Marco van Leeuwen ; Dr. Anton Machacek ; Dr. Komal Maheshwari ; Kosma von Maltitz ; Nicole Masureik ; John Mathew ; JoEllen McBride ; Nikolai Meures ; Riana Meyer ; Jenny Miller ; Abdul Mirza ; Asogan Moodaly ; Jothi Moodley ; Nolene Naidu ; Tyrone Negus ; Thomas O'Donnell ; Dr. Markus Oldenburg ; Dr. Jaynie Padayachee ; Nicolette Pekeur ; Sirika Pillay ; Jacques Plaut ; Andrea Prinsloo ; Joseph Raimondo ; Sanya Rajani ; Prof. Sergey Rakityansky ; Alastair Ramlakan ; Razvan Remsing ; Max Richter ; Sean Riddle ; Evan Robinson ; Dr. Andrew Rose ; Bianca Ruddy ; Katie Russell ; Duncan Scott ; Helen Seals ; Ian Sherratt ; Roger Sieloff ; Bradley Smith ; Greg Solomon ; Mike Stringer ; Shen Tian ; Robert Torregrosa ; Jimmy Tseng ; Helen Waugh ; Dr. Dawn Webber ; Michelle Wen ; Dr. Alexander Wetzler ; Dr. Spencer Wheaton ; Vivian White ; Dr. Gerald Wigger ; Harry Wiggins ; Wendy Williams ; Julie Wilson ; Andrew Wood ; Emma Wormauld ; Sahal Yacoob ; Jean Youssef

Contributors and editors have made a sincere effort to produce an accurate and useful resource. Should you have suggestions, find mistakes or be prepared to donate material for inclusion, please don't hesitate to contact us. We intend to work with all who are willing to help make this a continuously evolving resource!

## www.fhsst.org

# Contents

I	Int	roduction	1
11	M	atter and Materials	3
1	Clas	sification of Matter - Grade 10	5
	1.1	Mixtures	5
		1.1.1 Heterogeneous mixtures	6
		1.1.2 Homogeneous mixtures	6
		1.1.3 Separating mixtures	7
	1.2	Pure Substances: Elements and Compounds	9
		1.2.1 Elements	9
		1.2.2 Compounds	9
	1.3	Giving names and formulae to substances	10
	1.4	Metals, Semi-metals and Non-metals	13
		1.4.1 Metals	13
		1.4.2 Non-metals	14
		1.4.3 Semi-metals	14
	1.5	Electrical conductors, semi-conductors and insulators	14
	1.6	Thermal Conductors and Insulators	15
	1.7	Magnetic and Non-magnetic Materials	17
	1.8	Summary	18
2	Wha	at are the objects around us made of? - Grade 10	21
	2.1	Introduction: The atom as the building block of matter	21
	2.2	Molecules	21
		2.2.1 Representing molecules	21
	2.3	Intramolecular and intermolecular forces	25
	2.4	The Kinetic Theory of Matter	26
	2.5	The Properties of Matter	28
	2.6	· Summary	31
2	<b>-</b> .		25
3			35
	3.1	Models of the Atom	35
		3.1.1 The Plum Pudding Model	35
		3.1.2 Rutherford's model of the atom	36

		3.1.3 The Bohr Model	87
	3.2	How big is an atom?	88
		3.2.1 How heavy is an atom?	88
		3.2.2 How big is an atom?	88
	3.3	Atomic structure	8
		3.3.1 The Electron	39
		3.3.2 The Nucleus	39
	3.4	Atomic number and atomic mass number	0
	3.5	lsotopes	2
		3.5.1 What is an isotope?	2
		3.5.2 Relative atomic mass	5
	3.6	Energy quantisation and electron configuration	6
		3.6.1 The energy of electrons	6
		3.6.2 Energy quantisation and line emission spectra 4	7
		3.6.3 Electron configuration	7
		3.6.4 Core and valence electrons	51
		3.6.5 The importance of understanding electron configuration 5	51
	3.7	Ionisation Energy and the Periodic Table	3
		3.7.1 lons	53
		3.7.2 Ionisation Energy	5
	3.8	The Amongsment of Atoms in the Davidia Table	6
	3.0	The Arrangement of Atoms in the Periodic Table 5	0
	3.0		6
	3.0	3.8.1 Groups in the periodic table	
	3.9	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5	6
4	3.9	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5	6 8
4	3.9	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5	6 8 9 <b>3</b>
4	3.9 Ator	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6	6 8 9 <b>3</b>
4	3.9 <b>Ator</b> 4.1	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6	56 58 59 <b>3</b>
4	<ul><li>3.9</li><li>Ator</li><li>4.1</li><li>4.2</li></ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6	56 58 59 <b>3</b> 53
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6	56 58 59 53 53 55
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         4.4.1 The nature of the covalent bond       6	i6 i8 i9 <b>3</b> i3 i5 i5
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         4.4.1 The nature of the covalent bond       6         Lewis notation and molecular structure       6	i6 i8 i9 <b>3</b> i3 i5 i5 i5
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5 <b>ic Combinations - Grade 11</b> 6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         4.4.1 The nature of the covalent bond       6         Lewis notation and molecular structure       6	i6 i8 i9 <b>3</b> i3 i5 i5 i5 i9 2
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         4.4.1 The nature of the covalent bond       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1 Non-polar and polar covalent bonds       7	i6 i8 i9 <b>3</b> i3 i5 i5 i5 i9 2
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> </ul>	3.8.1       Groups in the periodic table       5         3.8.2       Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1       Non-polar and polar covalent bonds       7         4.6.2       Polar molecules       7	i6 i8 i9 <b>3</b> i3 i5 i5 i5 i5 i2 i3 i3 i5 i5 i5 i2 i3 i3 i3 i5 i5 i5 i5 i5 i5 i5 i5 i5 i5 i5 i5 i5
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> <li>4.6</li> </ul>	3.8.1 Groups in the periodic table       5         3.8.2 Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1 Non-polar and polar covalent bonds       7         4.6.2 Polar molecules       7         Ionic Bonding       7	i6 i8 i9 <b>3</b> i3 i3 i5 i5 i9 2 3 3
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> <li>4.6</li> </ul>	3.8.1       Groups in the periodic table       5         3.8.2       Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1       Non-polar and polar covalent bonds       7         4.6.2       Polar molecules       7         Ionic Bonding       7         4.7.1       The nature of the ionic bond       7	i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4         i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4         i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4         i7       i7       i8       i7       i7
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> <li>4.6</li> </ul>	3.8.1       Groups in the periodic table       5         3.8.2       Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1       Non-polar and polar covalent bonds       7         4.6.2       Polar molecules       7         Ionic Bonding       7         4.7.1       The nature of the ionic bond       7         4.7.2       The crystal lattice structure of ionic compounds       7	i6         i8         i9         i3         i3         i5         i5         i5         i6         i8         i9         i3         i3         i5
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> <li>4.6</li> </ul>	3.8.1       Groups in the periodic table       5         3.8.2       Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         Version of the covalent bond       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1       Non-polar and polar covalent bonds       7         4.6.2       Polar molecules       7         Ionic Bonding       7         4.7.1       The nature of the ionic bond       7         4.7.2       The crystal lattice structure of ionic compounds       7         4.7.3       Properties of lonic Compounds       7	i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4       i4         i6       i8       i9       i3       i3       i5       i5       i5       i9       i2       i3       i4       i4       i6
4	<ul> <li>3.9</li> <li>Ator</li> <li>4.1</li> <li>4.2</li> <li>4.3</li> <li>4.4</li> <li>4.5</li> <li>4.6</li> <li>4.7</li> </ul>	3.8.1       Groups in the periodic table       5         3.8.2       Periods in the periodic table       5         Summary       5         sic Combinations - Grade 11       6         Why do atoms bond?       6         Energy and bonding       6         What happens when atoms bond?       6         Covalent Bonding       6         4.4.1       The nature of the covalent bond       6         Lewis notation and molecular structure       6         Electronegativity       7         4.6.1       Non-polar and polar covalent bonds       7         4.6.2       Polar molecules       7         Ionic Bonding       7         4.7.1       The nature of the ionic bond       7         4.7.2       The crystal lattice structure of ionic compounds       7         4.7.3       Properties of lonic Compounds       7         Metallic bonds       7	i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4       i4       i6         i6       i8       i9       i3       i3       i5       i5       i5       i9       i2       i3       i4       i4       i6       i6         i6       i8       i9       i3       i3       i5       i5       i9       i2       i3       i4       i4       i6       i6         i7       i8       i8       i9       i8       i8

	4.9	Writing chemical formulae
		4.9.1 The formulae of covalent compounds
		4.9.2 The formulae of ionic compounds $\ldots \ldots 80$
	4.10	The Shape of Molecules
		4.10.1 Valence Shell Electron Pair Repulsion (VSEPR) theory $\hdots\hdo$
		4.10.2 Determining the shape of a molecule
	4.11	Oxidation numbers
	4.12	Summary
5	Inte	rmolecular Forces - Grade 11 91
	5.1	Types of Intermolecular Forces
	5.2	$Understanding\ intermolecular\ forces \qquad \ldots \qquad \ldots \qquad \ldots \qquad . \qquad . \qquad . \qquad . \qquad . \qquad . \qquad .$
	5.3	Intermolecular forces in liquids
	5.4	Summary
6	Solu	tions and solubility - Grade 11 101
	6.1	Types of solutions
	6.2	Forces and solutions
	6.3	Solubility
	6.4	Summary
7	Ator	nic Nuclei - Grade 11 107
	7.1	Nuclear structure and stability
	7.2	The Discovery of Radiation
	7.3	Radioactivity and Types of Radiation
		7.3.1 Alpha ( $lpha$ ) particles and alpha decay
		7.3.2 Beta ( $\beta$ ) particles and beta decay $\ldots \ldots 109$
		7.3.3 Gamma ( $\gamma$ ) rays and gamma decay
	7.4	Sources of radiation
		7.4.1 Natural background radiation
		7.4.2 Man-made sources of radiation
	7.5	The 'half-life' of an element
	7.6	The Dangers of Radiation
	7.7	The Uses of Radiation
	7.8	Nuclear Fission
		7.8.1 The Atomic bomb - an abuse of nuclear fission
		7.8.2 Nuclear power - harnessing energy
	7.9	Nuclear Fusion
	7.10	Nucleosynthesis
		7.10.1 Age of Nucleosynthesis (225 s - $10^3$ s)
		7.10.2 Age of lons $(10^3 \text{ s} - 10^{13} \text{ s})$
		7.10.3 Age of Atoms $(10^{13} \text{ s} - 10^{15} \text{ s})$
		/
		7.10.3 Age of Atoms (10-5 S - 10-5 S)       122         7.10.4 Age of Stars and Galaxies (the universe today)       122

8	Ther	rmal Properties and Ideal Gases - Grade 11 1	25
	8.1	A review of the kinetic theory of matter $\ldots$ $\ldots$ $\ldots$ $\ldots$ $\ldots$ $1$	.25
	8.2	Boyle's Law: Pressure and volume of an enclosed gas	.26
	8.3	Charles's Law: Volume and Temperature of an enclosed gas	.32
	8.4	The relationship between temperature and pressure $\ldots \ldots \ldots$	.36
	8.5	The general gas equation	.37
	8.6	The ideal gas equation	.40
	8.7	Molar volume of gases	.45
	8.8	Ideal gases and non-ideal gas behaviour	.46
	8.9	Summary	.47
9	Orga	anic Molecules - Grade 12 1	51
	9.1	What is organic chemistry?	.51
	9.2	Sources of carbon	.51
	9.3	Unique properties of carbon	.52
	9.4	Representing organic compounds	.52
		9.4.1 Molecular formula	.52
		9.4.2 Structural formula	.53
		9.4.3 Condensed structural formula	.53
	9.5	Isomerism in organic compounds	.54
	9.6	Functional groups	.55
	9.7	The Hydrocarbons	.55
		9.7.1 The Alkanes	.58
		9.7.2 Naming the alkanes	.59
		9.7.3 Properties of the alkanes	.63
		9.7.4 Reactions of the alkanes	.63
		9.7.5 The alkenes	.66
		9.7.6 Naming the alkenes	.66
		9.7.7 The properties of the alkenes	.69
		9.7.8 Reactions of the alkenes	.69
		9.7.9 The Alkynes	.71
		9.7.10 Naming the alkynes	.71
	9.8	The Alcohols	.72
		9.8.1 Naming the alcohols	.73
		9.8.2 Physical and chemical properties of the alcohols	.75
	9.9	Carboxylic Acids	.76
		9.9.1 Physical Properties	.77
		9.9.2 Derivatives of carboxylic acids: The esters	.78
	9.10	The Amino Group	.78
	9.11	The Carbonyl Group	.78
	9.12	Summary	.79

10	Orga	nnic Macromolecules - Grade 12	185
	10.1	Polymers	185
	10.2	How do polymers form?	186
		10.2.1 Addition polymerisation	186
		10.2.2 Condensation polymerisation	188
	10.3	The chemical properties of polymers	190
	10.4	Types of polymers	191
	10.5	Plastics	191
		10.5.1 The uses of plastics	192
		10.5.2 Thermoplastics and thermosetting plastics	194
		10.5.3 Plastics and the environment	195
	10.6	Biological Macromolecules	196
		10.6.1 Carbohydrates	197
		10.6.2 Proteins	199
		10.6.3 Nucleic Acids	202
	10.7	Summary	204
			200
	C	hemical Change	209
11	Phys	sical and Chemical Change - Grade 10	211
	11.1	Physical changes in matter	211
	11.2	Chemical Changes in Matter	212
		11.2.1 Decomposition reactions	213
		11.2.2 Synthesis reactions	214
	11.3	Energy changes in chemical reactions	217
	11.4	Conservation of atoms and mass in reactions	217
	11.5	Law of constant composition	219
	11.6	Volume relationships in gases	219
	11.7	Summary	220
12	Repr	resenting Chemical Change - Grade 10	223
	12.1	Chemical symbols	223
	12.2	Writing chemical formulae	224
	12.3	Balancing chemical equations	224
		12.3.1 The law of conservation of mass	224
		12.3.2 Steps to balance a chemical equation	226
	12.4	State symbols and other information	230
	12.5	Summary	232
13	Qua	ntitative Aspects of Chemical Change - Grade 11	233

13.1	The Mole	33
13.2	Molar Mass	35
13.3	An equation to calculate moles and mass in chemical reactions	37

13.4	${\sf Molecules \ and \ compounds}  \ldots  \ldots  \ldots  \ldots  \ldots  \ldots  \ldots  \ldots  \ldots  $
13.5	The Composition of Substances
13.6	Molar Volumes of Gases
13.7	Molar concentrations in liquids
13.8	Stoichiometric calculations
13.9	Summary
11 Eno	rgy Changes In Chemical Reactions - Grade 11 255
	What causes the energy changes in chemical reactions?
	Exothermic and endothermic reactions
	The heat of reaction
	Examples of endothermic and exothermic reactions
	Spontaneous and non-spontaneous reactions
	Activation energy and the activated complex
	Summary
14.7	Summary
15 Тур	es of Reactions - Grade 11 267
15.1	Acid-base reactions
	15.1.1 What are acids and bases?
	15.1.2 Defining acids and bases
	15.1.3 Conjugate acid-base pairs
	15.1.4 Acid-base reactions
	15.1.5 Acid-carbonate reactions
15.2	Redox reactions
	15.2.1 Oxidation and reduction
	15.2.2 Redox reactions
15.3	Addition, substitution and elimination reactions
	15.3.1 Addition reactions
	15.3.2 Elimination reactions
	15.3.3 Substitution reactions
15.4	Summary
16 Dec	ation Dates Crada 12 207
	ction Rates - Grade 12 287
	Introduction
	Factors affecting reaction rates    289
	Reaction rates and collision theory
	Measuring Rates of Reaction
	Mechanism of reaction and catalysis
16.6	Chemical equilibrium
	16.6.1 Open and closed systems
	16.6.2 Reversible reactions
	16.6.3 Chemical equilibrium
16.7	The equilibrium constant

353

		16.7.1	Calculating the equilibrium constant	5
		16.7.2	The meaning of $k_c$ values	5
	16.8	Le Cha	telier's principle	С
		16.8.1	The effect of concentration on equilibrium	С
		16.8.2	The effect of temperature on equilibrium	С
		16.8.3	The effect of pressure on equilibrium	2
	16.9	Industr	ial applications	5
	16.10	)Summa	ary	5
17	Elect	rochen	nical Reactions - Grade 12 319	9
	17.1	Introdu	ction	9
	17.2	The Ga	Ilvanic Cell	)
		17.2.1	Half-cell reactions in the Zn-Cu cell $\ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots 32$	1
		17.2.2	Components of the Zn-Cu cell	2
		17.2.3	The Galvanic cell	3
		17.2.4	Uses and applications of the galvanic cell $\ldots \ldots \ldots \ldots \ldots \ldots \ldots 324$	4
	17.3	The Ele	ectrolytic cell	5
		17.3.1	The electrolysis of copper sulphate	5
		17.3.2	The electrolysis of water	7
		17.3.3	A comparison of galvanic and electrolytic cells	3
	17.4	Standa	rd Electrode Potentials	3
		17.4.1	The different reactivities of metals	9
		17.4.2	Equilibrium reactions in half cells	9
		17.4.3	Measuring electrode potential	С
		17.4.4	The standard hydrogen electrode	С
		17.4.5	Standard electrode potentials	3
		17.4.6	Combining half cells	7
		17.4.7	Uses of standard electrode potential	3
	17.5	Balanci	ing redox reactions	2
	17.6	Applica	tions of electrochemistry	7
		17.6.1	Electroplating	7
		17.6.2	The production of chlorine	3
		17.6.3	Extraction of aluminium	9
	17.7	Summa	ary	9

# **IV** Chemical Systems

18	The Water Cycle - Grade 10	355
	18.1 Introduction	355
	18.2 The importance of water	355
	18.3 The movement of water through the water cycle	356
	18.4 The microscopic structure of water	359

18.4.1 The polar nature of water	. 359
18.4.2 Hydrogen bonding in water molecules	. 359
18.5 The unique properties of water	. 360
18.6 Water conservation	. 363
18.7 Summary	. 366
19 Global Cycles: The Nitrogen Cycle - Grade 10	369
19.1 Introduction	. 369
19.2 Nitrogen fixation	. 369
19.3 Nitrification	. 371
19.4 Denitrification	. 372
19.5 Human Influences on the Nitrogen Cycle	. 372
19.6 The industrial fixation of nitrogen	. 373
19.7 Summary	. 374
20 The Hydrosphere - Grade 10	377
20.1 Introduction	. 377
20.2 Interactions of the hydrosphere	. 377
20.3 Exploring the Hydrosphere	. 378
20.4 The Importance of the Hydrosphere	. 379
20.5 lons in aqueous solution	. 379
20.5.1 Dissociation in water	. 380
20.5.2 lons and water hardness	. 382
20.5.3 The pH scale	. 382
20.5.4 Acid rain	. 384
20.6 Electrolytes, ionisation and conductivity	. 386
20.6.1 Electrolytes	. 386
20.6.2 Non-electrolytes	. 387
20.6.3 Factors that affect the conductivity of water	. 387
20.7 Precipitation reactions	. 389
20.8 Testing for common anions in solution	. 391
20.8.1 Test for a chloride	. 391
20.8.2 Test for a sulphate	. 391
20.8.3 Test for a carbonate	. 392
20.8.4 Test for bromides and iodides	. 392
20.9 Threats to the Hydrosphere	. 393
20.10Summary	. 394
21 The Lithosphere - Grade 11	397
21.1 Introduction	. 397
21.2 The chemistry of the earth's crust	. 398
21.3 A brief history of mineral use	. 399
21.4 Energy resources and their uses	. 400

21.5 Mining and Mineral Processin	g: Gold
21.5.1 Introduction	
21.5.2 Mining the Gold	
21.5.3 Processing the gold or	e
21.5.4 Characteristics and us	es of gold
21.5.5 Environmental impact	s of gold mining 404
21.6 Mining and mineral processing	g: Iron
21.6.1 Iron mining and iron c	re processing
21.6.2 Types of iron	
21.6.3 Iron in South Africa .	
21.7 Mining and mineral processing	g: Phosphates
21.7.1 Mining phosphates .	
21.7.2 Uses of phosphates .	
21.8 Energy resources and their use	es: Coal
21.8.1 The formation of coal	
21.8.2 How coal is removed f	rom the ground
21.8.3 The uses of coal	
21.8.4 Coal and the South A	rican economy
21.8.5 The environmental im	pacts of coal mining
21.9 Energy resources and their use	es: Oil
21.9.1 How oil is formed $\ldots$	
21.9.2 Extracting oil	
21.9.3 Other oil products	
21.9.4 The environmental im	pacts of oil extraction and use
21.10Alternative energy resources .	
21.11Summary	
22 The Atmosphere - Grade 11	421
22.1 The composition of the atmos	phere
22.2 The structure of the atmosph	ere
22.2.1 The troposphere	
22.2.2 The stratosphere	
22.2.3 The mesosphere	
22.2.4 The thermosphere	
22.3 Greenhouse gases and global	varming
22.3.1 The heating of the atr	nosphere
22.3.2 The greenhouse gases	and global warming
22.3.3 The consequences of g	lobal warming
22.3.4 Taking action to comb	bat global warming
22.4 Summary	

23 The Chemical Industry - Grade 12	435
23.1 Introduction	435
23.2 Sasol	435
23.2.1 Sasol today: Technology and production	436
23.2.2 Sasol and the environment	440
23.3 The Chloralkali Industry	442
23.3.1 The Industrial Production of Chlorine and Sodium Hydroxide	442
23.3.2 Soaps and Detergents	446
23.4 The Fertiliser Industry	450
23.4.1 The value of nutrients	450
23.4.2 The Role of fertilisers	450
23.4.3 The Industrial Production of Fertilisers	451
23.4.4 Fertilisers and the Environment: Eutrophication	454
23.5 Electrochemistry and batteries	456
23.5.1 How batteries work	456
23.5.2 Battery capacity and energy	457
23.5.3 Lead-acid batteries	457
23.5.4 The zinc-carbon dry cell	459
23.5.5 Environmental considerations	460
23.6 Summary	461

A GNU Free Documentation License

467

# Chapter 7

# Atomic Nuclei - Grade 11

**Nuclear physics** is the branch of physics which deals with the **nucleus** of the atom. Within this field, some scientists focus their attention on looking at the *particles* inside the nucleus and understanding how they interact, while others classify and interpret the *properties* of nuclei. This detailed knowledge of the nucleus makes it possible for *technological advances* to be made. In this next chapter, we are going to touch on each of these different areas within the field of nuclear physics.

## 7.1 Nuclear structure and stability

You will remember from an earlier chapter that an atom is made up of different types of particles: protons (positive charge) neutrons (neutral) and electrons (negative charge). The nucleus is the part of the atom that contains the protons and the neutrons, while the electrons are found in energy orbitals around the nucleus. The protons and neutrons together are called **nucleons**. It is the nucleus that makes up most of an atom's *atomic mass*, because an electron has a very small mass when compared with a proton or a neutron.

Within the nucleus, there are different forces which act between the particles. The **strong nuclear force** is the force between two or more nucleons, and this force binds protons and neutrons together inside the nucleus. This force is most powerful when the nucleus is small, and the nucleons are close together. The **electromagnetic force** causes the repulsion between like-charged (positive) protons. In a way then, these forces are trying to produce opposite effects in the nucleus. The strong nuclear force acts to hold all the protons and neutrons close together, while the electromagnetic force acts to push protons further apart. In atoms where the nuclei are small, the strong nuclear force overpowers the electromagnetic force. However, as the nucleus gets bigger (in elements with a higher number of nucleons), the electromagnetic force becomes greater than the strong nuclear force. In these nuclei, it becomes possible for particles and energy to be ejected from the nucleus. These nuclei are called **unstable**. The particles and energy that a nucleus releases are referred to as **radiation**, and the atom is said to be **radioactive**. We are going to look at these concepts in more detail in the next few sections.

## 7.2 The Discovery of Radiation

Radioactivity was first discovered in 1896 by a French scientist called Henri Becquerel while he was working on phosphorescent materials. He happened to place some uranium crystals on black paper that he had used to cover a piece of film. When he looked more carefully, he noticed that the film had lots of patches on it, and that this did not happen when other elements were placed on the paper. He eventually concluded that some rays must be coming out of the uranium crystals to produce this effect.

His observations were taken further by the Polish scientist Marie Curie and her husband Pierre, who increased our knowledge of radioactive elements. In 1903, Henri, Marie and Pierre were

awarded the Nobel Prize in Physics for their work on radioactive elements. This award made Marie the first woman ever to receive a Nobel Prize. Marie Curie and her husband went on to discover two new elements, which they named **polonium** (Po) after Marie's home country, and **radium** (Ra) after its highly radioactive characteristics. For these dicoveries, Marie was awarded a Nobel Prize in Chemistry in 1911, making her one of very few people to receive two Nobel Prizes.



7.3

Marie Curie died in 1934 from aplastic anemia, which was almost certainly partly due to her massive exposure to radiation during her lifetime. Most of her work was carried out in a shed without safety measures, and she was known to carry test tubes full of radioactive isotopes in her pockets and to store them in her desk drawers. By the end of her life, not only was she very ill, but her hands had become badly deformed due to their constant exposure to radiation. Unfortunately it was only later in her life that the full dangers of radiation were realised.

## 7.3 Radioactivity and Types of Radiation

In section 7.1, we discussed that when a nucleus is unstable it can emit particles and energy. This is called **radioactive decay**.



#### **Definition: Radioactive decay**

Radioactive decay is the process in which an unstable atomic nucleus loses energy by emitting particles or electromagnetic waves. **Radiation** is the name for the emitted particles or electromagnetic waves.

When a nucleus undergoes radioactive decay, it emits radiation and the nucleus is called radioactive. We are exposed to small amounts of radiation all the time. Even the rocks around us emit radiation! However some elements are far more radioactive than others. *Isotopes* tend to be less stable because they contain a larger number of nucleons than 'non-isotopes' of the same element. These radioactive isotopes are called **radioisotopes**.

Radiation can be emitted in different forms. There are three main types of radiation: alpha, beta and gamma radiation. These are shown in figure 7.1, and are described below.

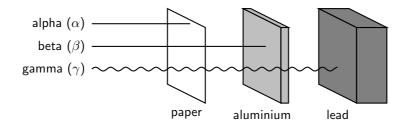


Figure 7.1: Types of radiation

### 7.3.1 Alpha ( $\alpha$ ) particles and alpha decay

An alpha particle is made up of two protons and two neutrons bound together. This type of radiation has a *positive charge*. An alpha particle is sometimes represented using the chemical symbol  $He^{2+}$ , because it has the same structure as a Helium atom (two neutrons and two protons) which is missing its two electrons, hence the overall charge of +2. Alpha particles have very low penetration power. Penetration power describes how easily the particles can pass through another material. Because alpha particles have a *low* penetration power, it means that even something as thin as a piece of paper or the outside layer of the human skin, will absorb these particles so that they can't go any further.

Alpha decay occurs because the nucleus has too many protons, and this causes a lot of repulsion between these like charges. To try to reduce this repulsion, the nucleus emits an  $\alpha$  particle. This can be seen in the decay of Americium (Am) to Neptunium (Np).

Example:

$$^{241}_{95}$$
Am  $\rightarrow ^{237}_{93}$  Np +  $\alpha$  particle

Let's take a closer look at what has happened during this reaction. Americium (Z = 95; A = 241) undergoes  $\alpha$  decay and releases one alpha particle (i.e. 2 protons and 2 neutrons). The atom now has only 93 protons (Z = 93). On the periodic table, the element which has 93 protons (Z = 93) is called Neptunium. Therefore, the Americium atom has become a Neptunium atom. The atomic mass of the neptunium atom is 237 (A = 237) because 4 nucleons (2 protons and 2 neutrons) were emitted from the atom of Americium.

## **7.3.2** Beta ( $\beta$ ) particles and beta decay

In certain types of radioactive nuclei that have too many neutrons, a neutron may be converted into a proton, an electron and another particle (called a *neutrino*). The high energy electrons that are released in this way are called **beta particles**. Beta particles have a higher penetration power than alpha particles and are able to pass through thicker materials such as paper.

The diagram below shows what happens during  $\beta$  decay:

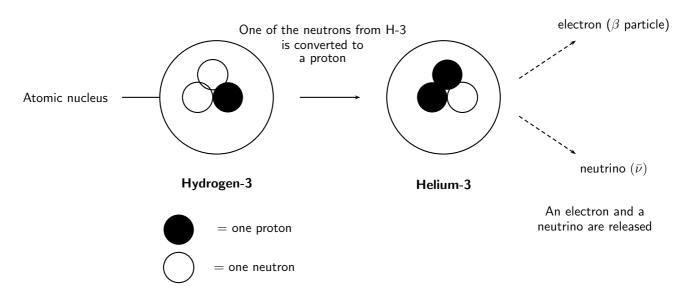


Figure 7.2:  $\beta$  decay in a hydrogen atom

During beta decay, the number of neutrons in the atom decreases by one, and the number of protons increases by one. Since the number of protons before and after the decay is different,

the atom has changed into a different element. In figure 7.2, Hydrogen has become Helium. The beta decay of the Hydrogen-3 atom can be represented as follows:

 $^{3}_{1}\text{H} \rightarrow ^{3}_{2}\text{He} + \beta \text{particle} + \bar{\nu}$ 



When scientists added up all the energy from the neutrons, protons and electrons involved in  $\beta$ -decays, they noticed that there was always some energy missing. We know that energy is always conserved, which led Wolfgang Pauli in 1930 to come up with the idea that another particle, which was not detected yet, also had to be involved in the decay. He called this particle the neutrino (Italian for "little neutral one"), because he knew it had to be neutral, have little or no mass, and interact only very weakly, making it very hard to find experimentally! The neutrino was finally identified experimentally about 25 years after Pauli first thought of it.

Due to the radioactive processes inside the sun, each 1  $cm^2$  patch of the earth receives 70 billion  $(70 \times 10^9)$  neutrinos each second! Luckily neutrinos only interact very weakly so they do not harm our bodies when billions of them pass through us every second.

#### 7.3.3 Gamma ( $\gamma$ ) rays and gamma decay

When particles inside the nucleus collide during radioactive decay, energy is released. This energy can leave the nucleus in the form of waves of electromagnetic energy called gamma rays. Gamma radiation is part of the electromagnetic spectrum, just like visible light. However, unlike visible light, humans cannot see gamma rays because they are at a higher frequency and a higher energy. Gamma radiation has no mass or charge. This type of radiation is able to penetrate most common substances, including metals. The only substances that can absorb this radiation are thick lead and concrete.

Gamma decay occurs if the nucleus is at too high an energy level. Since gamma rays are part of the electromagnetic spectrum, they can be thought of as waves or particles. Therefore in gamma decay, we can think of a ray or a particle (called a photon) being released. The atomic number and atomic mass remain unchanged.

Table 7.1 summarises and compares the three types of radioactive decay that have been discussed.

Type of decay	Particle/ray released	Change in element	Penetration power
Alpha ( $\alpha$ )	lpha particle (2 protons and 2 neutrons)	Yes	Low
Beta ( $\beta$ )	$\beta$ particle (electron)	Yes	Medium
Gamma ( $\gamma$ )	$\gamma$ ray (electromagnetic energy)	No	High



#### Worked Example 23: Radioactive decay

Question: The isotope  ${}^{241}_{95}$ Pb undergoes radioactive decay and loses two alpha particles.

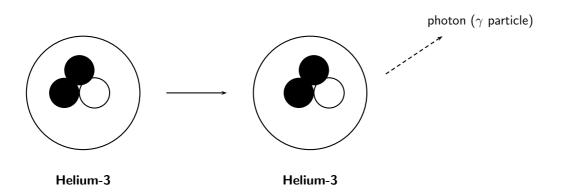


Figure 7.3:  $\gamma$  decay in a helium atom

- 1. Write the chemical formula of the element that is produced as a result of the decay.
- 2. Write an equation for this decay process.

#### Answer

#### Step 1 : Work out the number of protons and/or neutrons that the radioisotope loses during radioactive decay

One  $\alpha$  particle consists of two protons and two neutrons. Since two  $\alpha$  particles are released, the total number of protons lost is four and the total number of neutrons lost is also four.

Step 2 : Calculate the atomic number (Z) and atomic mass number (A) of the element that is formed.

$$Z = 95 - 4 = 91$$
  
 $A = 241 - 4 = 237$ 

Step 3 : Refer to the periodic table to see which element has the atomic number that you have calculated.

The element that has Z = 91 is Protactinium (Pa).

Step 4 : Write the symbol for the element that has formed as a result of radioactive decay.

 $^{237}_{91}$ Pa

Step 5 : Write an equation for the decay process.

 $^{241}_{95}Pb \rightarrow ^{237}_{91}Pa + 2$  protons + 2 neutrons

#### Activity :: Discussion : Radiation

In groups of 3-4, discuss the following questions:

• Which of the three types of radiation is most dangerous to living creatures (including humans!)

- What can happen to people if they are exposed to high levels of radiation?
- What can be done to protect yourself from radiation (Hint: Think of what the radiologist does when you go for an X-ray)?

?

#### Exercise: Radiation and radioactive elements

- 1. There are two main forces inside an atomic nucleus:
  - (a) Name these two forces.
  - (b) Explain why atoms that contain a greater number of nucleons are more likely to be radioactive.
- 2. The isotope  $\frac{241}{95}$ Pb undergoes radioactive decay and loses three alpha particles.
  - (a) Write the chemical formula of the element that is produced as a result of the decay.
  - (b) How many nucleons does this element contain?
- 3. Complete the following equation:

 $^{210}_{82}Am \rightarrow (alpha decay)$ 

- 4. Radium-228 decays by emitting a beta particle. Write an equation for this decay process.
- 5. Describe how gamma decay differs from alpha and beta decay.

## 7.4 Sources of radiation

The sources of radiation can be either natural or man-made.

## 7.4.1 Natural background radiation

Cosmic radiation

The Earth, and all living things on it, are constantly bombarded by radiation from space. Charged particles from the sun and stars interact with the Earth's atmosphere and magnetic field to produce a shower of radiation, mostly beta and gamma radiation. The amount of cosmic radiation varies in different parts of the world because of differences in elevation and also the effects of the Earth's magnetic field.

• Terrestrial Radiation

Radioactive material is found throughout nature. It occurs naturally in the soil, water, and vegetation. The major isotopes that are of concern are uranium and the decay products of uranium, such as thorium, radium, and radon. Low levels of uranium, thorium, and their decay products are found everywhere. Some of these materials are ingested (taken in) with food and water, while others are breathed in. The dose of radiation from terrestrial sources varies in different parts of the world.



Cosmic and terrestrial radiation are not the only natural sources. All people have radioactive potassium-40, carbon-14, lead-210 and other isotopes inside their bodies from birth.

### 7.4.2 Man-made sources of radiation

Although all living things are exposed to natural background radiation, there are other sources of radiation. Some of these will affect most members of the public, while others will only affect those people who are exposed to radiation through their work.

#### • Members of the Public

Man-made radiation sources that affect members of the public include televisions, tobacco (polonium-210), combustible fuels, smoke detectors (americium), luminous watches (tritium) and building materials. By far, the most significant source of man-made radiation exposure to the public is from medical procedures, such as diagnostic x-rays, nuclear medicine, and radiation therapy. Some of the major isotopes involved are I-131, Tc-99m, Co-60, Ir-192, and Cs-137. The production of nuclear fuel using uranium is also a source of radiation for the public, as is fallout from nuclear weapons testing or use.

• Individuals who are exposed through their work

Any people who work in the following environments are exposed to radiation at some time: radiology (X-ray) departments, nuclear power plants, nuclear medicine departments and radiation oncology (the study of cancer) departments. Some of the isotopes that are of concern are cobalt-60, cesium-137, americium-241, and others.



Radiation therapy (or radiotherapy) uses ionising radiation as part of cancer treatment to control malignant cells. In cancer, a malignant cell is one that divides very rapidly to produce many more cells. These groups of dividing cells can form a growth or **tumour**. The malignant cells in the tumour can take nutrition away from other healthy body cells, causing them to die, or can increase the pressure in parts of the body because of the space that they take up. Radiation therapy uses radiation to try to target these malignant cells and kill them. However, the radiation can also damage other, healthy cells in the body. To stop this from happening, shaped radiation beams are aimed from several angles to intersect at the tumour, so that the radiation dose here is much higher than in the surrounding, healthy tissue. But even doing this doesn't protect all the healthy cells, and that is why people have side-effects to this treatment.

Note that radiation therapy is different from chemotherapy, which uses *chemicals*, rather than radiation, to destroy malignant cells. Generally, the side effects of chemotherapy are greater because the treatment is not as localised as it is with radiation therapy. The chemicals travel throughout the body, affecting many healthy cells.

## 7.5 The 'half-life' of an element



#### Definition: Half-life

The half-life of an element is the time it takes for half the atoms of a radioisotope to decay into other atoms.

Radioisotope	Chemical symbol	Half-life	
Polonium-212	Po-212	0.16 seconds	
Sodium-24	Na-24	15 hours	
Strontium-90	Sr-90	28 days	
Cobalt-60	Co-60	5.3 years	
Caesium-137	Cs-137	30 years	
Carbon-14	C-14	5 760 years	
Calcium	Ca	100 000 years	
Beryllium	Be	2 700 000 years	
Uranium-235	U-235	7.1 billion years	

Table 7.2: Table showing the half-life of a number of elements

So, in the case of Sr-90, it will take 28 days for half of the atoms to decay into other atoms. It will take another 28 days for half of the remaining atoms to decay. Let's assume that we have a sample of strontium that weighs 8g. After the first 28 days there will be:

$$1/2 \times 8 = 4$$
 g Sr-90 left

After 56 days, there will be:

$$1/2 \ge 4 \ g = 2 \ g \ Sr-90 \ left$$

After 84 days, there will be:

$$1/2 \ge 2 = 1 = 1 = 1$$
 s Sr-90 left

If we convert these amounts to a *fraction* of the original sample, then after 28 days 1/2 of the sample remains undecayed. After 56 days 1/4 is undecayed and after 84 days, 1/8 and so on.

Activity :: Group work : Understanding half-life Work in groups of 4-5 You will need: 16 sheets of A4 paper per group, scissors, 2 boxes per group, a marking pen and timer/stopwatch.

#### What to do:

- Your group should have two boxes. Label one 'decayed' and the other 'radioactive'.
- Take the A4 pages and cut each into 4 pieces of the same size. You should now have 64 pieces of paper. Stack these neatly and place them in the 'radioactive' box. The paper is going to represent some radioactive material.
- Set the timer for one minute. After one minute, remove half the sheets of paper from the radioactive box and put them in the 'decayed' box.
- Set the timer for another minute and repeat the previous step, again removing half the pieces of paper that are left in the radioactive box and putting them in the decayed box.
- Repeat this process until 8 minutes have passed. You may need to start cutting your pieces of paper into even smaller pieces as you progress.

#### Questions:

- 1. How many pages were left in the radioactive box after...
  - (a) 1 minute
  - (b) 3 minutes

- (c) 5 minutes
- 2. What percentage (%) of the pages were left in the radioactive box after...
  - (a) 2 minutes
  - (b) 4 minutes
- 3. After how many minutes is there 1/128 of radioactive material remaining?
- 4. What is the half-life of the 'radioactive' material in this exercise?



### Worked Example 24: Half-life 1

**Question:** A 100 g sample of Cs-137 is allowed to decay. Calculate the mass of Cs-137 that will be left after 90 years

#### Answer

**Step 1 : You need to know the half-life of Cs-137** The half-life of Cs-137 is 30 years.

# Step 2 : Determine how many times the quantity of sample will be halved in 90 years.

If the half-life of Cs-137 is 30 years, and the sample is left to decay for 90 years, then the number of times the quantity of sample will be halved is 90/30 = 3.

# Step 3 : Calculate the quantity that will be left by halving the mass of Cs-137 three times

1. After 30 years, the mass left is 100 g  $\times$  1/2 = 50 g

2. After 60 years, the mass left is 50 g  $\times$  1/2 = 25 g

3. After 90 years, the mass left is 25 g  $\times$  1/2 = 12.5 g

Note that a quicker way to do this calculation is as follows: Mass left after 90 years =  $(1/2)^3 \times 100$  g = 12.5 g (The exponent is the number of times the quantity is halved)



#### Worked Example 25: Half-life 2

**Question:** An 80 g sample of Po-212 decays until only 10 g is left. How long did it take for this decay to take place?

#### Answer

Step 1 : Calculate the fraction of the original sample that is left after decay Fraction remaining =10~g/80~g=1/8

Step 2 : Calculate how many half-life periods of decay (x) must have taken place for 1/8 of the original sample to be left

$$(\frac{1}{2})^x = \frac{1}{8}$$

#### Therefore, x = 3

7.6

Step 3 : Use the half-life of Po-212 to calculate how long the sample was left to decay

The half-life of Po-212 is 0.16 seconds. Therefore if there were three periods of decay, then the total time is 0.16  $\times$  3. The time that the sample was left to decay is 0.48 seconds.

#### Exercise: Looking at half life

- 1. Imagine that you have 100 g of Na-24.
  - (a) What is the half life of Na-24?
  - (b) How much of this isotope will be left after 45 hours?
  - (c) What percentage of the original sample will be left after 60 hours?
- 2. A sample of Sr-90 is allowed to decay. After 84 days, 10 g of the sample remains.
  - (a) What is the half life of Sr-90?
  - (b) How much Sr-90 was in the original sample?
  - (c) How much Sr-90 will be left after 112 days?

## 7.6 The Dangers of Radiation

Natural radiation comes from a variety of sources such as the rocks, sun and from space. However, when we are exposed to large amounts of radiation, this can cause damage to cells.  $\gamma$ radiation is particularly dangerous because it is able to penetrate the body, unlike  $\alpha$  and  $\beta$ particles whose penetration power is less. Some of the dangers of radiation are listed below:

#### • Damage to cells

Radiation is able to penetrate the body, and also to penetrate the membranes of the cells within our bodies, causing massive damage. *Radiation poisoning* occurs when a person is exposed to large amounts of this type of radiation. Radiation poisoning damages tissues within the body, causing symptoms such as diarrhoea, vomiting, loss of hair and convulsions.

#### • Genetic abnormalities

When radiation penetrates cell membranes, it can damage chromosomes within the nucleus of the cell. The chromosomes contain all the genetic information for that person. If the chromosomes are changed, this may lead to genetic abnormalities in any children that are born to the person who has been exposed to radiation. Long after the nuclear disaster of Chernobyl in Russia in 1986, babies were born with defects such as missing limbs and abnormal growths.

#### • Cancer

Small amounts of radiation can cause cancers such as leukemia (cancer of the blood)

## 7.7 The Uses of Radiation

However, despite the many dangers of radiation, it does have many powerful uses, some of which are listed below:

#### • Medical Field

Radioactive *chemical tracers* emitting  $\gamma$  rays can give information about a person's internal anatomy and the functioning of specific organs. The radioactive material may be injected into the patient, from where it will target specific areas such as bones or tumours. As the material decays and releases radiation, this can be seen using a special type of camera or other instrument. The radioactive material that is used for this purpose must have a short half-life so that the radiation can be detected quickly and also so that the material is quickly removed from the patient's body. Using radioactive materials for this purpose can mean that a tumour or cancer may be diagnosed long before these would have been detected using other methods such as X-rays.

Radiation may also be used to sterilise medical equipment.

Activity :: Research Project : The medical uses of radioisotopes

Carry out your own research to find out more about the radioisotopes that are used to diagnose diseases in the following parts of the body:

- thyroid gland
- kidneys
- brain

In each case, try to find out...

- 1. which radioisotope is used
- 2. what the sources of this radioisotope are
- 3. how the radioisotope enters the patient's body and how it is monitored

#### • Biochemistry and Genetics

Radioisotopes may be used as tracers to label molecules so that chemical processes such as DNA replication or amino acid transport can be traced.

#### • Food preservation

Irradiation of food can stop vegetables or plants from sprouting after they have been harvested. It also kills bacteria and parasites, and controls the ripening of fruits.

#### Environment

Radioisotopes can be used to trace and analyse pollutants.

#### • Archaeology and Carbon dating

Natural radioisotopes such as C-14 can be used to determine the age of organic remains. All living organisms (e.g. trees, humans) contain carbon. Carbon is taken in by plants and trees through the process of photosynthesis in the form of carbon dioxide and is then converted into organic molecules. When animals feed on plants, they also obtain carbon through these organic compounds. Some of the carbon in carbon dioxide is the radioactive C-14, while the rest is a non-radioactive form of carbon. When an organism dies, no more carbon is taken in and there is a limited amount of C-14 in the body. From this point onwards, C-14 begins its radioactive decay. When scientists uncover remains, they are able to estimate the age of the remains by seeing how much C-14 is left in the body relative to the amount of non-radioactive carbon. The less C-14 there is, the older the remains because radioactive decay must have been taking place for a long time. Because scientists know the exact rate of decay of C-14, they can calculate a very accurate estimate of the age of the remains. Carbon dating has been a very important tool in building up accurate historical records.

#### Activity :: Case Study : Using radiocarbon dating

Radiocarbon dating has played an important role in uncovering many aspects of South Africa's history. Read the following extract from an article that appeared in Afrol news on 10th February 2007 and then answer the questions that follow.

The world famous rock art in South Africa's uKhahlamba-Drakensberg, a World Heritage Site, is three times older than previously thought, archaeologists conclude in a new study. The more than 40,000 paintings were made by the San people some 3000 years ago, a new analysis had shown.

Previous work on the age of the rock art in uKhahlamba-Drakensberg concluded it is less than 1,000 years old. But the new study - headed by a South African archaeologist leading a team from the University of Newcastle upon Tyne (UK) and Australian National University in Canberra - estimates the panels were created up to 3,000 years ago. They used the latest radio-carbon dating technology.

The findings, published in the current edition of the academic journal 'South African Humanities', have "major implications for our understanding of how the rock artists lived and the social changes that were taking place over the last three millennia," according to a press release from the British university.

#### Questions:

- 1. What is the half-life of carbon-14?
- 2. In the news article, what role did radiocarbon dating play in increasing our knowledge of South Africa's history?
- 3. Radiocarbon dating can also be used to analyse the remains of once-living organisms. Imagine that a set of bones are found between layers of sediment and rock in a remote area. A group of archaeologists carries out a series of tests to try to estimate the age of the bones. They calculate that the bones are approximately 23 040 years old.

What percentage of the original carbon-14 must have been left in the bones for them to arrive at this estimate?

## 7.8 Nuclear Fission

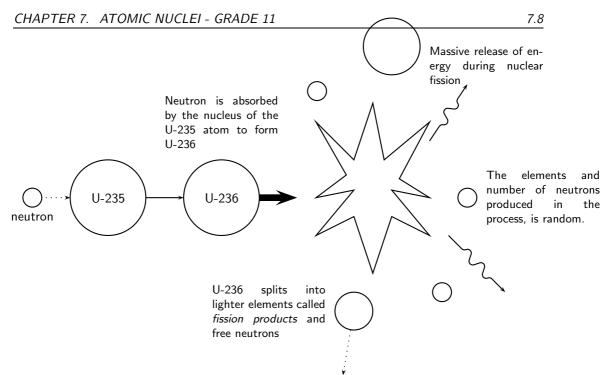
**Nuclear fission** is a process where the nucleus of an atom is split into two or more smaller nuclei, known as *fission products*. The fission of heavy elements is an **exothermic reaction** and huge amounts of energy are released in the process. This energy can be used to produce *nuclear power* or to make *nuclear weapons*, both of which we will discuss a little later.



Definition: Nuclear fission

The splitting of an atomic nucleus

Below is a diagram showing the nuclear fission of Uranium-235. An atom of Uranium-235 is bombarded with a neutron to initiate the fission process. This neutron is absorbed by Uranium-235, to become Uranium-236. Uranium-236 is highly unstable and breaks down into a number of lighter elements, releasing energy in the process. Free neutrons are also produced during this process, and these are then available to bombard other fissionable elements. This process is known as a **fission chain reaction**, and occurs when one nuclear reaction starts off another, which then also starts off another one so that there is a rapid increase in the number of nuclear reactions that are taking place.



## 7.8.1 The Atomic bomb - an abuse of nuclear fission

A nuclear chain reaction can happen very quickly, releasing vast amounts of energy in the process. In 1939, it was discovered that Uranium could undergo nuclear fission. In fact, it was uranium that was used in the first atomic bomb. The bomb contained huge amounts of Uranium-235, enough to start a runaway nuclear fission chain reaction. Because the process was uncontrolled, the energy from the fission reactions was released in a matter of *seconds*, resulting in the massive explosion of that first bomb. Since then, more atomic bombs have been dropped, causing massive destruction and loss of life.

Activity :: Discussion : Nuclear weapons testing - an ongoing issue Read the article below which has been adapted from one that appeared in 'The Globe' in Washington on 10th October 2006, and then answer the questions that follow.

US officials and arms control specialists warned yesterday that North Korea's test of a small nuclear device could start an arms race in the region and threaten the landmark global treaty designed nearly four decades ago to halt the spread of nuclear weapons. US officials expressed concern that North Korea's neighbors, including Japan, Taiwan, and South Korea, could eventually decide to develop weapons of their own. They also fear that North Korea's moves could embolden Iran, and that this in turn could encourage Saudi Arabia or other neighbours in the volatile Middle East to one day seek nuclear deterrents, analysts say.

North Korea is the first country to conduct a nuclear test after pulling out of the Nuclear Nonproliferation Treaty. The treaty, which was created in 1968, now includes 185 nations (nearly every country in the world). Under the treaty, the five declared nuclear powers at the time (United States, the Soviet Union, France, China, and Great Britain) agreed to reduce their supplies of nuclear weapons. The treaty has also helped to limit the number of new nuclear weapons nations.

But there have also been serious setbacks. India and Pakistan, which never signed the treaty, became new nuclear powers, shocking the world with test explosions in 1998. The current issue of nuclear weapons testing in North Korea, is another such setback and a blow to the treaty.

#### Group discussion questions:

1. Discuss what is meant by an 'arms race' and a 'treaty'.

- Do you think it is important to have such treaties in place to control the testing and use of nuclear weapons? Explain your answer.
- 3. Discuss some of the reasons why countries might not agree to be part of a nuclear weapons treaty.
- 4. How would you feel if South Africa decided to develop its own nuclear weapons?

## 7.8.2 Nuclear power - harnessing energy

However, nuclear fission can also be carried out in a controlled way in a *nuclear reactor*. A nuclear reactor is a piece of equpiment where nuclear chain reactions can be started in a controlled and sustained way. This is different from a nuclear *explosion* where the chain reaction occurs in seconds. The most important use of nuclear reactors at the moment is to produce **electrical power**, and most of these nuclear reactors use nuclear fission. A **nuclear fuel** is a chemical isotope that can keep a fission chain reaction going. The most common isotopes that are used are Uranium-235 and Plutonium-239. The amount of free energy that is in nuclear fuels is far greater than the energy in a similar amount of other fuels such as gasoline. In many countries, nuclear power is seen as a relatively environmentally friendly alternative to fossil fuels, which release large amounts of greenhouse gases, and are also non-renewable resources. However, one of the concerns around the use of nuclear power, is the production of *nuclear waste* which contains radioactive chemical elements.

#### Activity :: Debate : Nuclear Power

The use of nuclear power as a source of energy has been a subject of much debate. There are many advantages of nuclear power over other energy sources. These include the large amount of energy that can be produced at a small plant, little atmospheric pollution and the small quantity of waste. However there are also disadvantages. These include the expense of maintaining nuclear power stations, the huge impact that an accident could have as well as the disposal of dangerous nuclear waste.

Use these ideas as a starting point for a class debate.

#### Nuclear power - An energy alternative or environmental hazard?

Your teacher will divide the class into teams. Some of the teams will be 'pro' nuclear power while the others will be 'anti' nuclear power.

## 7.9 Nuclear Fusion

**Nuclear fusion** is the joining together of the nuclei of two atoms to form a heavier nucleus. If the atoms involved are small, this process is accompanied by the release of energy. It is the nuclear fusion of elements that causes stars to shine and hydrogen bombs to explode. As with nuclear *fission* then, there are both positive and negative uses of nuclear fusion.



The joining together of the nuclei of two atoms.

You will remember that nuclei naturally repel one another because of the electrostatic force between their positively charged protons. So, in order to bring two nuclei together, a lot of energy must be supplied if fusion is to take place. If two nuclei can be brought close enough together however, the electrostatic force is overwhelmed by the more powerful strong nuclear force which only operates over short distances. If this happens, nuclear fusion can take place. Inside the cores of stars, the temperature is high enough for hydrogen fusion to take place but scientists have so far been unsuccessful in making it work in the laboratory. One of the huge advantages of nuclear fusion, if it could be made to happen, is that it is a relatively environmentally friendly source of energy. The helium that is produced is not radioactive or poisonous and does not carry the dangers of nuclear fission.

## 7.10 Nucleosynthesis

An astronomer named Edwin Hubble discovered in the 1920's that the universe is expanding. He measured that far-away galaxies are moving away from the earth at great speed, and the further away they are, the faster they are moving.



#### Extension: What are galaxies?

Galaxies are huge clusters of stars and matter in the universe. The earth is part of the Milky Way galaxy which is shaped like a very large spiral. Astronomers can measure the light coming from distant galaxies using telescopes. Edwin Hubble was also able to measure the velocities of galaxies.

These observations led people to see that the universe is expanding. It also led to the *Big Bang* hypothesis. The 'Big Bang' hypothesis is an idea about how the universe may have started. According to this theory, the universe started off at the beginning of time as a point which then exploded and expanded into the universe we live in today. This happened between 10 and 14 billion years ago.

Just after the Big Bang, when the universe was only  $10^{-43}$ s old, it was very hot and was made up of quarks and leptons (an example of a lepton is the electron). As the universe expanded, ( $\sim 10^{-2}$ s) and cooled, the quarks started binding together to form protons and neutrons (together called *nucleons*).

## **7.10.1** Age of Nucleosynthesis (225 s - $10^3$ s)

About 225 s after the Big Bang, the protons and neutrons started binding together to form simple *nuclei*. The process of forming nuclei is called *nucleosynthesis*. When a proton and a neutron bind together, they form the *deuteron*. The deuteron is like a hydrogen nucleus (which is just a proton) with a neutron added to it so it can be written as <sup>2</sup>H. Using protons and neutrons as building blocks, more nuclei can be formed as shown below. For example, the Helium-4 nucleus (also called an *alpha particle*) can be formed in the following ways:

 $^{2}H + n \rightarrow ^{3}H$ deuteron + neutron  $\rightarrow$  triton

then:

$${}^{3}\text{H} + p \rightarrow {}^{4}\text{He}$$
  
triton + proton  $\rightarrow$  Helium4 (alpha particle)

 $^{2}H + p \rightarrow ^{3}He$ deuteron + proton  $\rightarrow$  Helium3 *then:* 

 $^{3}\text{He} + n \rightarrow ^{4}\text{He}$ Helium3 + neutron  $\rightarrow$  Helium4 (alpha particle)

Some  $^{7}Li$  nuclei could also have been formed by the fusion of  $^{4}He$  and  $^{3}H$ .

## **7.10.2** Age of lons ( $10^3$ s - $10^{13}$ s)

However, at this time the universe was still very hot and the electrons still had too much energy to become bound to the alpha particles to form helium *atoms*. Also, the nuclei with mass numbers greater than 4 (i.e. greater than <sup>4</sup>He) are very short-lived and would have decayed almost immediately after being formed. Therefore, the universe moved through a stage called the Age of lons when it consisted of free positively charged  $\rm H^+$  ions and <sup>4</sup>He ions, and negatively charged electrons not yet bound into atoms.

## **7.10.3** Age of Atoms ( $10^{13}$ s - $10^{15}$ s)

As the universe expanded further, it cooled down until the electrons were able to bind to the hydrogen and helium nuclei to form hydrogen and helium atoms. Earlier, during the Age of lons, both the hydrogen and helium ions were positively charged which meant that they repelled each other (electrostatically). During the Age of Atoms, the hydrogen and helium along with the electrons, were in the form of atoms which are electrically neutral and so they no longer repelled each other and instead pulled together under gravity to form clouds of gas, which evetually formed stars.

### 7.10.4 Age of Stars and Galaxies (the universe today)

Inside the core of stars, the densities and temperatures are high enough for fusion reactions to occur. Most of the heavier nuclei that exist today were formed inside stars from thermonuclear reactions! (It's interesting to think that the atoms that we are made of were actually manufactured inside stars!). Since stars are mostly composed of hydrogen, the first stage of thermonuclear reactions inside stars involves hydrogen and is called **hydrogen burning**. The process has three steps and results in four hydrogen atoms being formed into a helium atom with (among other things) two photons (light!) being released.

The next stage is **helium burning** which results in the formation of carbon. All these reactions release a large amount of energy and heat the star which causes heavier and heavier nuclei to fuse into nuclei with higher and higher atomic numbers. The process stops with the formation of <sup>56</sup>Fe, which is the most strongly bound nucleus. To make heavier nuclei, even higher energies are needed than is possible inside normal stars. These nuclei are most likely formed when huge amounts of energy are released, for example when stars explode (an exploding star is called a **supernova**). This is also how all the nuclei formed inside stars get "recycled" in the universe to become part of new stars and planets.

## 7.11 Summary

• Nuclear physics is the branch of physics that deals with the nucleus of an atom.

- There are two forces between the particles of the nucleus. The **strong nuclear force** is an attractive force between the neutrons and the **electromagnetic force** is the repulsive force between like-charged protons.
- In atoms with large nuclei, the electromagnetic force becomes greater than the strong nuclear force and particles or energy may be released from the nucleus.
- Radioactive decay occurs when an unstable atomic nucleus loses energy by emitting particles or electromagnetic waves.
- The particles and energy released are called **radiation** and the atom is said to be **radioactive**.
- Radioactive isotopes are called radioisotopes.
- Radioactivity was first discovered by Marie Curie and her husband Pierre.
- There are three types of radiation from radioactive decay: **alpha** ( $\alpha$ ), **beta** ( $\beta$ ) and **gamma** ( $\gamma$ ) radiation.
- During **alpha decay**, an alpha particle is released. An alpha particle consists of two protons and two neutrons bound together. Alpha radiation has low penetration power.
- During **beta decay**, a beta particle is released. During beta decay, a neutron is converted to a proton, an electron and a neutrino. A beta particle is the electron that is released. Beta radiation has greater penetration power than alpha radiation.
- During **gamma decay**, electromagnetic energy is released as gamma rays. Gamma radiation has the highest penetration power of the three radiation types.
- There are many sources of radiation. Some of natural and others are man-made.
- Natural sources of radiation include cosmic and terrestrial radiation.
- Man-made sources of radiation include televisions, smoke detectors, X-rays and radiation therapy.
- The **half-life** of an element is the time it takes for half the atoms of a radioisotope to decay into other atoms.
- Radiation can be very damaging. Some of the negative impacts of radiation exposure include damage to cells, genetic abnormalities and cancer.
- However, radiation can also have many **positive uses**. These include use in the medical field (e.g. chemical tracers), biochemistry and genetics, use in food preservation, the environment and in archaeology.
- Nuclear fission is the splitting of an atomic nucleus into smaller fission products. Nuclear fission produces large amounts of energy, which can be used to produce nuclear power, and to make nuclear weapons.
- **Nuclear fusion** is the joining together of the nuclei of two atoms to form a heavier nucleus. In stars, fusion reactions involve the joining of hydrogen atoms to form helium atoms.
- **Nucleosynthesis** is the process of forming nuclei. This was very important in helping to form the universe as we know it.



### Exercise: Summary exercise

- 1. Explain each of the following terms:
  - (a) electromagnetic force
  - (b) radioactive decay
  - (c) radiocarbon dating

- 2. For each of the following questions, choose the **one correct answer**:
  - (a) The part of the atom that undergoes radioactive decay is the...
    - i. neutrons
    - ii. nucleus
    - iii. electrons
    - iv. entire atom
  - (b) The radio-isotope Po-212 undergoes alpha decay. Which of the following statements is **true**?
    - i. The number of protons in the element remains unchanged.
    - ii. The number of nucleons after decay is 212.
    - iii. The number of protons in the element after decay is 82.
    - iv. The end product after decay is Po-208.
- 3. 20 g of sodium-24 undergoes radoactive decay. Calculate the percentage of the original sample that remains after 60 hours.
- 4. Nuclear physics can be controversial. Many people argue that studying the nucleus has led to devastation and huge loss of life. Others would argue that the benefits of nuclear physics far outweigh the negative things that have come from it.
  - (a) Outline some of the ways in which nuclear physics has been used in negative ways.
  - (b) Outline some of the benefits that have come from nuclear physics.

you must enclose the copies in covers that carry, clearly and legibly, all these Cover Texts: Front-Cover Texts on the front cover, and Back-Cover Texts on the back cover. Both covers must also clearly and legibly identify you as the publisher of these copies. The front cover must present the full title with all words of the title equally prominent and visible. You may add other material on the covers in addition. Copying with changes limited to the covers, as long as they preserve the title of the Document and satisfy these conditions, can be treated as verbatim copying in other respects.

If the required texts for either cover are too voluminous to fit legibly, you should put the first ones listed (as many as fit reasonably) on the actual cover, and continue the rest onto adjacent pages.

If you publish or distribute Opaque copies of the Document numbering more than 100, you must either include a machine-readable Transparent copy along with each Opaque copy, or state in or with each Opaque copy a computer-network location from which the general network-using public has access to download using public-standard network protocols a complete Transparent copy of the Document, free of added material. If you use the latter option, you must take reasonably prudent steps, when you begin distribution of Opaque copies in quantity, to ensure that this Transparent copy will remain thus accessible at the stated location until at least one year after the last time you distribute an Opaque copy (directly or through your agents or retailers) of that edition to the public.

It is requested, but not required, that you contact the authors of the Document well before redistributing any large number of copies, to give them a chance to provide you with an updated version of the Document.

# MODIFICATIONS

You may copy and distribute a Modified Version of the Document under the conditions of sections A and A above, provided that you release the Modified Version under precisely this License, with the Modified Version filling the role of the Document, thus licensing distribution and modification of the Modified Version to whoever possesses a copy of it. In addition, you must do these things in the Modified Version:

- 1. Use in the Title Page (and on the covers, if any) a title distinct from that of the Document, and from those of previous versions (which should, if there were any, be listed in the History section of the Document). You may use the same title as a previous version if the original publisher of that version gives permission.
- 2. List on the Title Page, as authors, one or more persons or entities responsible for authorship of the modifications in the Modified Version, together with at least five of the principal authors of the Document (all of its principal authors, if it has fewer than five), unless they release you from this requirement.
- 3. State on the Title page the name of the publisher of the Modified Version, as the publisher.
- 4. Preserve all the copyright notices of the Document.
- 5. Add an appropriate copyright notice for your modifications adjacent to the other copyright notices.
- Include, immediately after the copyright notices, a license notice giving the public permission to use the Modified Version under the terms of this License, in the form shown in the Addendum below.
- 7. Preserve in that license notice the full lists of Invariant Sections and required Cover Texts given in the Document's license notice.
- 8. Include an unaltered copy of this License.
- 9. Preserve the section Entitled "History", Preserve its Title, and add to it an item stating at least the title, year, new authors, and publisher of the Modified Version as given on the Title Page. If there is no section Entitled "History" in the Document, create one stating the title, year, authors, and publisher of the Document as given on its Title Page, then add an item describing the Modified Version as stated in the previous sentence.

- 10. Preserve the network location, if any, given in the Document for public access to a Transparent copy of the Document, and likewise the network locations given in the Document for previous versions it was based on. These may be placed in the "History" section. You may omit a network location for a work that was published at least four years before the Document itself, or if the original publisher of the version it refers to gives permission.
- 11. For any section Entitled "Acknowledgements" or "Dedications", Preserve the Title of the section, and preserve in the section all the substance and tone of each of the contributor acknowledgements and/or dedications given therein.
- 12. Preserve all the Invariant Sections of the Document, unaltered in their text and in their titles. Section numbers or the equivalent are not considered part of the section titles.
- 13. Delete any section Entitled "Endorsements". Such a section may not be included in the Modified Version.
- 14. Do not re-title any existing section to be Entitled "Endorsements" or to conflict in title with any Invariant Section.
- 15. Preserve any Warranty Disclaimers.

If the Modified Version includes new front-matter sections or appendices that qualify as Secondary Sections and contain no material copied from the Document, you may at your option designate some or all of these sections as invariant. To do this, add their titles to the list of Invariant Sections in the Modified Version's license notice. These titles must be distinct from any other section titles.

You may add a section Entitled "Endorsements", provided it contains nothing but endorsements of your Modified Version by various parties-for example, statements of peer review or that the text has been approved by an organisation as the authoritative definition of a standard.

You may add a passage of up to five words as a Front-Cover Text, and a passage of up to 25 words as a Back-Cover Text, to the end of the list of Cover Texts in the Modified Version. Only one passage of Front-Cover Text and one of Back-Cover Text may be added by (or through arrangements made by) any one entity. If the Document already includes a cover text for the same cover, previously added by you or by arrangement made by the same entity you are acting on behalf of, you may not add another; but you may replace the old one, on explicit permission from the previous publisher that added the old one.

The author(s) and publisher(s) of the Document do not by this License give permission to use their names for publicity for or to assert or imply endorsement of any Modified Version.

# COMBINING DOCUMENTS

You may combine the Document with other documents released under this License, under the terms defined in section A above for modified versions, provided that you include in the combination all of the Invariant Sections of all of the original documents, unmodified, and list them all as Invariant Sections of your combined work in its license notice, and that you preserve all their Warranty Disclaimers.

The combined work need only contain one copy of this License, and multiple identical Invariant Sections may be replaced with a single copy. If there are multiple Invariant Sections with the same name but different contents, make the title of each such section unique by adding at the end of it, in parentheses, the name of the original author or publisher of that section if known, or else a unique number. Make the same adjustment to the section titles in the list of Invariant Sections in the license notice of the combined work.

In the combination, you must combine any sections Entitled "History" in the various original documents, forming one section Entitled "History"; likewise combine any sections Entitled "Ac-knowledgements", and any sections Entitled "Dedications". You must delete all sections Entitled "Endorsements".

# COLLECTIONS OF DOCUMENTS

You may make a collection consisting of the Document and other documents released under this License, and replace the individual copies of this License in the various documents with a single copy that is included in the collection, provided that you follow the rules of this License for verbatim copying of each of the documents in all other respects.

You may extract a single document from such a collection, and distribute it individually under this License, provided you insert a copy of this License into the extracted document, and follow this License in all other respects regarding verbatim copying of that document.

# AGGREGATION WITH INDEPENDENT WORKS

A compilation of the Document or its derivatives with other separate and independent documents or works, in or on a volume of a storage or distribution medium, is called an "aggregate" if the copyright resulting from the compilation is not used to limit the legal rights of the compilation's users beyond what the individual works permit. When the Document is included an aggregate, this License does not apply to the other works in the aggregate which are not themselves derivative works of the Document.

If the Cover Text requirement of section A is applicable to these copies of the Document, then if the Document is less than one half of the entire aggregate, the Document's Cover Texts may be placed on covers that bracket the Document within the aggregate, or the electronic equivalent of covers if the Document is in electronic form. Otherwise they must appear on printed covers that bracket the whole aggregate.

# TRANSLATION

Translation is considered a kind of modification, so you may distribute translations of the Document under the terms of section A. Replacing Invariant Sections with translations requires special permission from their copyright holders, but you may include translations of some or all Invariant Sections in addition to the original versions of these Invariant Sections. You may include a translation of this License, and all the license notices in the Document, and any Warranty Disclaimers, provided that you also include the original English version of this License and the original versions of those notices and disclaimers. In case of a disagreement between the translation and the original version of this License or a notice or disclaimer, the original version will prevail.

If a section in the Document is Entitled "Acknowledgements", "Dedications", or "History", the requirement (section A) to Preserve its Title (section A) will typically require changing the actual title.

# TERMINATION

You may not copy, modify, sub-license, or distribute the Document except as expressly provided for under this License. Any other attempt to copy, modify, sub-license or distribute the Document is void, and will automatically terminate your rights under this License. However, parties who have received copies, or rights, from you under this License will not have their licenses terminated so long as such parties remain in full compliance.

# FUTURE REVISIONS OF THIS LICENSE

The Free Software Foundation may publish new, revised versions of the GNU Free Documentation License from time to time. Such new versions will be similar in spirit to the present version, but may differ in detail to address new problems or concerns. See http://www.gnu.org/copyleft/.

Each version of the License is given a distinguishing version number. If the Document specifies that a particular numbered version of this License "or any later version" applies to it, you have the option of following the terms and conditions either of that specified version or of any later version that has been published (not as a draft) by the Free Software Foundation. If the Document does not specify a version number of this License, you may choose any version ever published (not as a draft) by the Free Software Foundation.

# ADDENDUM: How to use this License for your documents

To use this License in a document you have written, include a copy of the License in the document and put the following copyright and license notices just after the title page:

Copyright © YEAR YOUR NAME. Permission is granted to copy, distribute and/or modify this document under the terms of the GNU Free Documentation License, Version 1.2 or any later version published by the Free Software Foundation; with no Invariant Sections, no Front-Cover Texts, and no Back-Cover Texts. A copy of the license is included in the section entitled "GNU Free Documentation License".

If you have Invariant Sections, Front-Cover Texts and Back-Cover Texts, replace the "with...Texts." line with this:

with the Invariant Sections being LIST THEIR TITLES, with the Front-Cover Texts being LIST, and with the Back-Cover Texts being LIST.

If you have Invariant Sections without Cover Texts, or some other combination of the three, merge those two alternatives to suit the situation.

If your document contains nontrivial examples of program code, we recommend releasing these examples in parallel under your choice of free software license, such as the GNU General Public License, to permit their use in free software.